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Solving Rigid Body Problems

non-holonomic constraints \Rightarrow then use

Lagrange multipliers (rolling)
 holonomic constraints $f(\dot{\xi}, \ddot{\xi}) = 0$
 then choose independent $\dot{\xi}_j$

if axis of rotation is fixed, angles are easy
 if not fixed

- a) fixed point - use that for origin
- b) no fixed point - use CM for origin

use principal axes for coordinate system

$$\frac{d\vec{L}}{dt} = \underbrace{\left(\frac{d\vec{L}}{dt} \right)_\text{body}}_{\text{(body system)}} + \vec{\omega} \times \vec{L} = \vec{N}$$

$$L_j = I_j w_j \quad (\text{no summation})$$

~~Work~~. Work in body system (mostly)

$$\frac{dL_i}{dt} + \epsilon_{ijk} w_j L_k = N_i$$

$$\frac{dL_1}{dt} + w_2 L_3 - w_3 L_2 = N_1$$

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$$I_1 \dot{w}_1 - \beta w_2 w_3 (I_2 - I_3) = N_1$$

$$I_2 \dot{w}_2 - w_3 w_1 (I_3 - I_1) = N_2$$

$$I_3 \dot{w}_3 - w_1 w_2 (I_1 - I_2) = N_3$$

Cannot spin w/o Torque at constant \vec{w}
unless $\vec{w} \parallel \hat{n}_i$ (a principal axis)

Tire balancing: need principal axis \parallel to
axis of rotation

Specific case, Torque-free motion $\vec{N} = 0$
symmetric object $I_1 = I_2$

$$I_3 \ddot{w}_3 = 0 \quad \ddot{w}_1 + \beta w_3 w_2 = 0 \quad \beta = \frac{I_3 - I_1}{I_1}$$

$$\ddot{w}_2 - \beta w_2 w_1 = 0$$

$$\ddot{w}_1 + \beta w_3 \ddot{w}_2 = \ddot{w}_1 + \beta^2 w_3^2 w_1 = 0$$

$$\Rightarrow w_1 = A \cos(\beta w_3 t + \theta)$$

$$w_2 = A \sin(\beta w_3 t + \theta)$$

$$w = \sqrt{w_3^2 + A^2}$$

Precession! Instantaneous axis of rotation
traces a cone in the body system

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body cone $\frac{1}{2}$ angle α_b $A = w \sin \alpha_b$
 $\tan \alpha_b = A/w_3$ $w_3 = w \cos \alpha_b$

body is cotating around an axis that
is rotating

\vec{L} is constant

$$\cos \alpha_s = \frac{\vec{w} \cdot \vec{L}}{w L} = \frac{\vec{w}^\top \vec{I} \vec{w}}{w L} = \frac{ZT}{wL} \quad \text{space cone}$$

$$\begin{array}{lll} \alpha_b < \alpha_s & \text{PROLATE (CIGAR)} & I_3 > I_1 = I_2 \\ \alpha_b > \alpha_s & \text{OBBLATE (DISC)} & I_3 < I_1 = I_2 \end{array}$$

Nondegenerate $\overset{\leftrightarrow}{I}$ $I_1 \neq I_2 \neq I_3$
rotation is stable if $\vec{w} \parallel \hat{n}$.

perturbations?

a) $w_3 \gg w_1, w_2$ $w_3 \propto w, w_2$ 2nd order

$$w_1 = A [I_2 (I_3 - I_2)]^{1/2} \cos(\beta w_3 t + \theta)$$

$$w_2 = A [I_1 (I_3 - I_1)]^{1/2} \sin(\beta w_3 t + \theta)$$

$$\beta = \left[\frac{(I_3 - I_1)(I_3 - I_2)}{I_1 I_2} \right]^{1/2}$$

1) $I_3 > I_1, I_2 \Rightarrow \beta \in \mathbb{R}$ stable

2) $I_3 < I_1, I_2 \Rightarrow \beta \in \mathbb{R}$ stable

3) $I_1 < I_3 < I_2 \Rightarrow \beta \in \mathbb{C}$ unstable

tennis racket!

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More generally

$$I_1 \dot{w}_1 = w_2 w_3 (I_2 - I_3)$$

$$I_2 \dot{w}_2 = w_3 w_1 (I_3 - I_1)$$

$$I_3 \dot{w}_3 = w_1 w_2 (I_1 - I_2)$$

$$\vec{p} = \hat{n} / \sqrt{I} = \frac{\vec{w}}{w \sqrt{I}} = \frac{\vec{w}}{\sqrt{2T}}$$

$$T = \frac{\vec{w} \cdot \vec{L}}{2} = \frac{\vec{w}^T \overset{\leftrightarrow}{I} \vec{w}}{2} = \frac{w^2}{2} \hat{n}^T \overset{\leftrightarrow}{I} \hat{n} = \frac{1}{2} I w^2$$

$$\text{define } F(\vec{p}) = \vec{p}^T \overset{\leftrightarrow}{I} \vec{p} = p^2; I;$$

$F(\vec{p}) = 1$ defines the inertia ellipsoid

as \hat{n} and \vec{w} change, \vec{p} changes

But, the tip of \vec{p} stays on the inertia ellipsoid

$\vec{\nabla}_{\vec{p}} F$ is normal to the ellipsoid

$$\vec{\nabla}_{\vec{p}} F = 2 \overset{\leftrightarrow}{I} \vec{p} = \frac{2 \overset{\leftrightarrow}{I} \vec{w}}{\sqrt{2T}} = \sqrt{\frac{2}{T}} \vec{L}$$

\vec{w} constrained to move so that the
perp to the ellipsoid points \parallel to \vec{L}

\vec{L} is fixed in space

ellipsoid is fixed in the body

ellipsoid moves to keep the connection
between \vec{w} and \vec{L}

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distance from the origin of the ellipsoid
and the plane tangent to it @ \vec{P}

$$\frac{\vec{P} \cdot \vec{L}}{L} = \frac{\vec{w} \cdot \vec{L}}{L \sqrt{zT}} = \frac{\sqrt{zT}}{L} \text{ CONSTANT!}$$

$\vec{L} \perp$ invariable plane

ellipsoid rotates (rolls without slipping)
on the invariable plane

polhode = curve on the surface of the
ellipsoid

herpolhode = curve on the surface of
the invariable plane.

\vec{P} direction of \vec{w}
orientation of the ellipsoid \Rightarrow
orientation of the body

Case: Symmetric $I_1 = I_2$

inertia ellipsoid = ellipsoid of rotation

polhode = circle around the symmetry axis

herpolhode = circle

\vec{w} precesses around axis of symmetry

Now describe motion of \vec{L} wrt ⁶ body

$$T = \sum \frac{1}{2} \frac{\dot{L}_i^2}{I_i} = \text{constant}$$

$$\frac{L_1^2}{2TI_1} + \frac{L_2^2}{2TI_2} + \frac{L_3^2}{2TI_3} = 1 \quad \text{ellipsoid}$$

$$L_1^2 + L_2^2 + L_3^2 = L^2 \quad \text{sphere}$$

\vec{L} must lie on the sphere-ellipsoid intersection

$$\sqrt{2TI_3} \leq L \leq \sqrt{2TI_1} \quad (I_3 \leq I_2 \leq I_1)$$

Consider stability of motion

$L_x \quad L^2 = 2TI_1 + \epsilon \Rightarrow$ closed figures around L_x

$L_z \quad L^2 = 2TI_3 + \epsilon \Rightarrow$ closed figures around L_z

$L_y \quad$ no closed figures, large excursions! unstable

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Symmetric case $I_1 = I_2$

$$\omega_1 = A \cos(\beta w_3 + \theta) \quad \beta = \frac{I_3 - I_1}{I_1}$$

$$\omega_2 = A \sin(\beta w_3 + \theta)$$

Tidal bulge $\sim 30 \text{ km}$ at the equator
 $\sim 0.5\%$

$$\frac{I_3 - I_1}{I_1} \approx 3 \cdot 10^{-3}$$

period $\approx 300 \text{ days}$

expect precession around the axis with
 period $\approx 300 \text{ days}$

\Rightarrow deviations in "latitude" $\approx 10 \text{ m}$ (10^{-4})

annual variation: seasonal
 420-day period, due to precession?